RSC Advances



View Article Online

View Journal | View Issue

PAPER

Check for updates

Cite this: RSC Adv., 2023, 13, 12123

Received 13th February 2023 Accepted 6th April 2023

DOI: 10.1039/d3ra00978e

rsc.li/rsc-advances

1 Introduction

In the past few decades, various nanostructured materials have demonstrated outstanding potential in the energy field since the physicochemical properties of these materials can be engineered by replacing ions or doping with other elements with similar ionic radii and oxidation states.^{1,2} Furthermore, due to their versatile properties, such nanostructured materials can be effectively utilized in different fields for various applications. For example, in the photovoltaic area, thin-film technology based on nanostructured compound semiconductor materials offers numerous advantages over Si-based technology, as they exhibit enhanced optical properties, low production cost, and are environmentally friendly.^{1,3}

In recent times, semiconductor-based kesterite CZTS materials derived from CuInGaSe₂ (CIGS) by replacing Indium (In)

Enhanced photoresponse of Cu₂ZnSnS₄ absorber thin films fabricated using multi-metallic stacked nanolayers

Subhash Pandharkar,^{ad} Yogesh Hase,^a Shruti Shah,^a Vidya Doiphode,^a Ashish Waghmare,^a Ashvini Punde,^a Pratibha Shinde,^a Swati Rahane,^a Bharat Bade,^a Somnath Ladhane,^a Mohit Prasad,^{ab} Shashikant P. Patole^{10*c} and Sandesh Jadkar^{10*a}

Cu₂ZnSnS₄ (CZTS) thin films have attracted considerable attention as potential candidates for photovoltaic absorber materials. In a vacuum deposition technique, a sputtering stacked metallic layer followed by a thermal process for sulfur incorporation is used to obtain high-quality CZTS thin films. In this work, for fabricating CZTS thin films, we have done a 3LYS (3 layers), 6LYS, and 9LYS sequential deposition of Sn/ZnS/Cu metal stack (*via.* metallic stacked nanolayer precursors) onto Mo-coated corning glass substrate *via.* RF-sputtering. The prepared thin films were sulfurized in a tubular furnace at 550 °C in a gas mixture of 5% H₂S + 95% Ar for 10 min. We further investigated the impact of the Sn/ZnS/Cu metal stacking layers on the quality of the thin film based on its response to light because metal inter-diffusion during sulfurization is unavoidable. The inter-diffusion of precursors is low in a 3-layer stack sample, limiting the fabricated film's performance. CZTS films with 6-layer and 9-layer stacks result in an improved photocurrent density of ~19 μ A cm⁻². This enhancement can be attributed to the 9-layer approach's superior inter-diffusion of metallic precursors and compact, smooth CZTS microstructure evolution.

and Gallium (Ga) with earth-abundant Zinc (Zn) and Tin $(Sn)^{4-8}$ have been realized. Many investigations have been carried out in CZTS owing to its excellent properties. The constituent elements of CZTS are nontoxic and earth-abundant. It comprises a less hazardous substance, S, rather than Se, and has a direct bandgap energy of about 1.5 eV and a strong absorption coefficient $>10^4$ cm⁻¹ in the visible area.⁸⁻¹³ In recent years, by enhancing the synthesis and manufacturing procedure, CZTS-based solar cells' efficiency rose from 0.66% to 12.6%.¹⁴⁻¹⁶

Various deposition techniques, solution-based and vacuum process based,^{17–22} can be used to fabricate CZTS thin films. Solution-processed films behave like disordered systems, and controlling charge transport is always challenging in such films.²³ Such systems have significantly reduced mean free path and thereby decreased carrier mobilities than sputtered-based crystalline material.^{24,25} Sulfurizing CZT metallic precursor in a sulfur vapor (using pure H₂S gas) is essential for enhancing the crystal quality of the CZTS thin film. The sulfurization conditions and temperature determine the rate of crystallization reaction in the CZTS thin films.²⁶ The fabrication process for CZTS/CZTSe must adhere to a set of requirements, according to Scragg *et al.*²⁷ To prevent phase separation, all metal components and chalcogens are deposited simultaneously.

[&]quot;Department of Physics, Savitribai Phule Pune University, Pune, 411007, India. E-mail: sandesh@physics.unipune.ac.in

^bDepartment of Applied Science and Humanities, PCCOE, Nigdi, Pune, 411004, India ^cDepartment of Physics, Khalifa University of Science and Technology, Abu Dhabi, 127788, UAE. E-mail: shashikant.patole@ku.ac.ae

^dDepartment of Physics, Chandmal Tarachand Bora College, Shirur, Pune, 412210, India

However, we are far beyond the theoretically predicted efficiency, possibly owing to the small phase stability field, high series resistance, and Mo-CZTS interface recombination due to MoS₂ formation during sulfurization, Sn loss during heat treatment/sulfurization, and a range of potentially detrimental defects, etc.^{6,11,23,28,29} All these causes make it difficult to obtain a pure phase, controlled stoichiometry, and introduction of various secondary phases during the formation of CZTS thin films.^{25,30} Consequently, the enhancement in the photoresponse of kesterite absorbers is restricted.31 One of the challenges Yan et al.³² addressed was introducing an Al₂O₃ layer between the Mo electrode and absorber, which can serve as a protective layer preventing chalogenation of the Mo electrode, resulting in the reduction of Mo-chalcogenides transition layer and decrease in the series resistance of the device. Gour et al.33 observed that improved photocurrent is attributed to enhanced light harvesting, and it is influenced by the morphology of the material.34,35 Additionally, extra chalcogen vapor cab be supplied during annealing to prevent chalcogen and tin loss. The development of the precursor alloy, crystal growth, surface shape, and secondary phase with a particular distribution are all significantly influenced by the metallic precursor stacking sequence and result in various device performances.36 The Cu stack is typically the capping layer in a common stacking order of three metallic precursors, Sn/ZnS/Cu, because of its low volatility and high diffusion coefficient. Additionally, the Sn stack, which acts as the bottom layer, may be crucial for enhancing the dewetting effect, which improves thin film adhesion.37 However, although these ordinary synthesized three metallic precursors can suppress elemental loss, they cannot provide a homogeneous and uniform growth of films due to low diffusion of Sn in Cu (3.5 imes 10⁻³⁶ m² s⁻¹) and Zn in Cu (1.2 imes10⁻³⁸ m² s⁻¹) matrix.^{38,39}

It was noticed that the Zn to Cu matrix and the Sn to Zn matrix have shorter diffusion lengths than other matrices at a sulfurization temperature of 550 °C, as indicated in Table 1. With these shorter diffusion lengths, we can expect a better inter-diffusion in the 9LYS precursor during the sulfurization process, leading to a homogenous composition distribution in the resultant CZTS film.

In the present work, we have attempted to overcome the metal inter-diffusion problem using multi-stacked layers of metal precursors. We have prepared CZTS films on Mo-coated glass substrates using 3-layer metallic precursors in the sequence Sn/ZnS/Cu (3LYS). Additionally, using a repeating sequence of Sn/ZnS/Cu on glass substrates covered with Mo, we created 6-layer and 9-layer metallic precursors, referred to as 6-

Metal-to-metal matrix	Diffusion coefficient (a) 550 °C ($m^2 s^{-1}$)	Diffusion length in 10 min @ 550 °C (µm)					
Cu to Zn	$3.2 imes 10^{-12}$	44					
Cu to Sn	$1.9 imes10^{-9}$	1285					
Zn to Cu	2.6×10^{-17}	0.15					
Zn to Sn	7.2×10^{-10}	785					
Sn to Cu	9.7×10^{-17}	0.29					

Table 1 Inter diffusion coefficient in Cu/7n/Sn matrix³

LYS and 9LYS, respectively. The structural properties, morphological changes, optical properties, impurity phase formation, and photoresponse of 3LYS, 6LYS, and 9LYS samples are compared. We found that the photoresponse can be improved from ~19 μ A cm⁻² (*via* 3LYS) to ~38 μ A cm⁻² (*via* 6LYS) to ~82 μ A cm⁻² (*via* 9LYS) due to compact uniform microstructure and enhanced crystal quality along with suppression of secondary phase formation in 9LYS sample.

2 Experimental

2.1. Preparation of CZTS stacking layers

The CZTS thin films were prepared on Mo-coated glass substrates, followed by the sulfurization of deposited metallic precursors. The new procedure resulted in the precursor being progressively deposited from Cu (99.99%), ZnS (99.99%), and Sn (99.99%) targets at room temperature *via* RF sputtering, as shown in Fig. 1.

The stacking of the precursor's overall thickness was kept constant at 700 nm. We used 2 cm² Mo-coated corning glass substrates prepared by our own calibrated DC sputtering method⁴⁰ having a thickness of $\sim 1 \,\mu\text{m}$ and sheet resistance of 2 Ω per \Box . The Mo-coated glass substrates were cleaned using an ultrasonic cleaner for ten minutes each in acetone and deionized water. Then, they were blow-dried in N_2 (99.99%) gas to prepare the substrates for sputtering. Working gas pressure was maintained at 1.5 Pa (base pressure below 5×10^{-4} Pa), and precursors were kept in an atmosphere of Ar (99.99%). ZnS was deposited via sputtering power maintained at 120 W, and Cu and Sn required 70 W. Thickness of each precursor layer was controlled using deposition time. For example, the 3LYS metallic precursor was layered in the following order: Sn/ZnS/ Cu, which were then repeated twice for 6LYS and three times for 9LYS, all resulting in a thickness of \sim 700 nm. In the sulfurization process, all the multi-stacked precursor films are kept in a tubular furnace in an Ar atmosphere at ambient pressure. After that, heating was started, and at a temperature of 550 °C, 5% H₂S gas was mixed with Ar for the next 10 min. After 10 min, the H₂S supply was cut, and samples were rapidly cooled to room temperature in an Ar atmosphere.

2.2. Characterization of CZTS stacking layers

Various complementary characterization techniques have been used to analyze CZTS thin films. An X-ray diffractometer (XRD) (Bruker D8 Advance, Germany) was used to analyze the CZTS



Fig. 1 Deposition schematic of multi-metallic stacked-layer precursors.

Paper

thin films using the Cu-K line (=1.54) at a 1° grazing angle. The He-Ne (532 nm) laser was used to capture the Raman spectra by keeping laser power below 10 mW to prevent laser-induced crystallization. A JASCO UV-670 UV-visible spectrometer measured the optical absorption spectra between 400 and 1200 nm. An FEI, Nova NanoSem 450 scanning electron microscope was used to obtain field emission scanning electron micrographs (FE-SEM) of CZTS thin films. Thermo Scientific's K-Alpha+, UK machine with a resolution of 0.1 eV, was used to conduct X-ray photoelectron spectroscopy (XPS). Utilizing Al-K (1486.6 eV) radiation, we have captured XPS spectra for the particular element. After several scans in the acquisition process, the XPS signals were captured. By referring to C 1s at 284.6 eV, the binding energy was adjusted for specimen charging. A Metrohm Autolab: PGSTAT302N potentiostat and a Xenon lamp of 150 W (PEC-L01) assessed thin films' photocurrent density-time (I-t) characteristics.

3 Results and discussion

3.1. X-ray diffraction (XRD) analysis

As evident from Table 1, the different matrix has different diffusion lengths, and as expected, the 9LYS sample exhibited better inter-diffusion, resulting in homogenous composition and improved CZTS crystal quality. Therefore, XRD was used to confirm the phase formation and crystallinity of tetragonal kesterite-CZTS films. The diffractogram of these three sulfurized CZTS samples is shown in Fig. 2.

It should be noted that the Zn to Cu matrix and the Sn to Zn matrix have shorter diffusion lengths than other matrices at a sulfurization temperature of 550 °C, as indicated in Table 1.

As evident from Table 1, the different matrix has different diffusion lengths, and as expected, the 9LYS sample exhibited better inter-diffusion, resulting in homogenous composition in the resultant CZTS films. Therefore, X-RD was used to confirm the phase formation and crystallinity of tetragonal kesterite-



Fig. 2 XRD pattern of sputter-grown multi-stacked layers precursor CZTS films.

CZTS films. The diffractogram of these three sulfurized CZTS samples is shown in Fig. 2.

The prominent peak (112) of CZTS has different intensities, although there are no appreciable differences in the XRD peak positions. With more stacking layers, the (112) peak becomes more intense. Therefore, the 3LYS sample has weaker intensity compared to 6LYS and 9LYS samples. The overall volume of CZTS in a film can explain this decreased diffraction intensity. It should be noted that the Zn to Cu matrix and the Sn to Zn matrix have diffusion lengths of around 150 nm and 290 nm, respectively, at a sulfurization temperature of 550 °C (See Table 1). With these finite diffusion lengths, we anticipate superior inter-diffusion in the 9LYS precursor during the sulfurization process, resulting in a homogeneous composition and distribution in the resulting CZTS film. The prominent peaks are shown at 28.53°, 47.47°, and 56.29° correlate to the CZTS (112), (220), and (312) planes and are indicative of the kesterite structure (tetragonal system) [JCPDS card # 26-0575]. Some structural parameters like crystallite size (D), texture coefficient (TC), dislocation density (δ), average micro-strain (ϵ), etc., are extracted from the XRD pattern.

Scherrer's formula is used to determine the average crystallite size, $^{\!\!\!\!^{41}}$

$$D = \frac{K\lambda}{\beta\cos\theta_{\rm B}}\tag{1}$$

where *K* is the shape factor whose value is taken as 0.89, λ is the wavelength of the X-ray used, β is FWHM, and $\theta_{\rm B}$ is the Bragg diffraction angle.

The degree of preferred orientation of the different crystalline planes can be determined from Harris's analysis⁴² by calculating the texture coefficient (TC) using,

$$TC_{(hkl)} = \frac{I_{(hkl)}}{I_{o(hkl)}} \left[\frac{1}{n} \sum \frac{I_{(hkl)}}{I_{o(hkl)}} \right]^{-1}$$
(2)

The calculated values of the texture coefficient for stacking layers 3LYS, 6LYS, and 9LYS are listed in Table 2.

A crystallographic flaw or irregularity called a dislocation, which exists within a crystal structure, significantly impacts material properties. According to eqn (3), the dislocation density (δ) is the length of dislocation lines per unit volume of the crystal,⁴³

$$\delta = \frac{n}{D^2} \tag{3}$$

Table 2 Structural parameters, texture coefficient (TC), average crystalline size (*D*), lattice parameters (*a*, *b*, and *c*), dislocation density (δ), and average strain (ε) of stacked precursor CZTS thin films

Sample	TC (112)	D (nm)	$\delta imes 10^{-4}~{ m nm}^{-2}$	$\epsilon imes 10^{-3}$
3LYS	1.68	48	4.3	2.9
6LYS	1.73	41	5.9	3.4
9LYS	1.88	41	5.9	3.4

where D is the crystallite size and n is the factor that must equal one for the minimal dislocation density. Table 1 contains the calculated values of dislocation density.

RSC Advances

The displacement of the unit cell from its usual position causes the lattice micro-strain (ε) in a crystal. Lattice misfits, including flaws like dislocation, stacking fault likelihood, and lattice distortion, cause them to appear in CZTS thin films. Microstrain is calculated *via* the following equation,

$$\varepsilon = \frac{\beta \cos \theta_{\rm B}}{4} \tag{4}$$

The values of micro-strain for stacking layers 3LYS, 6LYS, and 9LYS are listed in Table 2. For the 9LYS CZTS film, the texture coefficient shows a much larger deviation from unity, which implies that the preferred orientation growth of the (112) plane is significantly higher than the 3LYS CZTS film. In addition, 9LYS CZTS film has a comparatively smaller crystallite size than 3LYS. However, it has a much higher crystallite volume fraction, which will help increase the charge carrier mobility.

3.2. Raman spectroscopy analysis

CZTS has a crystal structure identical to that of ZnS and Cu₂SnS₃, so it isn't easy to discern between the peaks of secondary phases (specifically ZnS, Cu₂SnS₃, and Cu₃SnS₄) and that of CZTS. Thus, fabricated films were investigated via. Raman spectroscopy is used to validate the presence of secondary phases, i. e. ZnS and Cu₂SnS₃. Fig. 3 shows Raman spectra for multi-stacked layer precursor 3LYS, 6LYS, and 9LYS CZTS thin films. The Raman spectra show a strong band at \sim 334 cm⁻¹, which agrees with the earlier reports on CZTS films.^{21,44} There were no additional peaks in the Raman spectra that corresponded to ZnS (350 cm⁻¹), Cu_{2-x}S (471-473 cm⁻¹), and Cu_2SnS_3 (298 cm⁻¹),⁴⁵⁻⁴⁷ indicating that the pure phase of kesterite-CZTS had formed. Chen et al.38 studied depth-resolved Raman spectroscopy of multi-stacked CZTS films. They also observed lower impurity levels and better metal inter-diffusion in the 9-layer stacking of precursor.



Fig. 3 Raman spectrum of multi-stacked layer 3LYS, 6LYS, and 9LYS CZTS films.



Fig. 4 XPS survey spectrum of CZTS 9LYS thin-film.

3.3. X-ray photoelectron spectroscopy analysis

Fig. 4 shows a typical survey spectrum of 9LYS CZTS thin film, which indicates that the presence of copper (Cu), zinc (Zn), tin (Sn), sulfur (S), and carbon (C) from reference and oxygen (O) from impurity. In Fig. 5(a) 9LYS sample shows pronounced splitting of the Cu 2p spectral line into the $2p_{1/2}$ and $2p_{3/2}$ core levels peaks at ~952.05 eV and ~ 932.2 eV with spectral line splitting of 19.85 eV that allows concluding that copper is in the +1 oxidation state indicating the formation of Cu(1).^{22,48} Fig. 5(b) 9LYS sample shows Zn 2p core level spectra which exhibit Zn $2p_{3/2}$ and $2p_{1/2}$ peaks visible at a binding energy of ~ 1045.2 and \sim 1022.05 eV with a spectral line separation of 23.15 eV suggesting the presence of zinc(II).^{22,48} The separation between Sn $3d_{5/2}$ and Sn $3d_{3/2}$ was 8.5 eV, shown in Fig. 5(c), indicating the presence of tin in the Sn(IV) state.22,48,49 The peak splitting between S 3p_{3/2} and 3p_{1/2} is 1.22 eV, which is in accordance with the expected S value (160-164 eV) in the sulfide phases.

3.4. Field-emission scanning electron microscopy analysis

Fig. 6 shows the surface morphology of sputter-grown 3LYS, 6LYS, and 9LYS multi-stacked layers of CZTS films. The 9-layer stacking precursor film showed well-defined CZTS grain and became well-faceted with densely packed granular film morphology [Fig. 6(c)]. Round particles packed on the surface were seen in the 3LYS precursor, and the 6LYS precursor showed rough morphology with discernible micro-voids [Fig. 6(c)], resulting in a poor interface.³⁶ The 9LYS precursor, meanwhile, yielded a thin, flat sheet that was relatively compact. Large agglomerated micrograins are observed in the 9LYS sample, which is advantageous from the perspective of solar cell applications since it lessens the recombination of photogenerated charge carriers. The average grain size of CZTS was calculated using the ImageJ application and significantly improved in the 9LYS precursor, which is \sim 0.7–0.8 µm than the conventional 3LYS sample.

The non-stoichiometry of compound semiconductors always correlates with their physicochemical properties. For example, for high-performing CZTS solar cell absorbers, there is an empirical rule that the Cu/(Zn + Sn) and Zn/Sn atomic ratios



Fig. 5 High-resolution core level XPS spectra of CZTS 9LYS thin-film (a) Cu-2p in the range 925–965 eV, (b) Zn-2p in the range 1015–1055 eV, (c) Sn-3d in the range 482–502 eV, and (d) S-2p in the range 156–168 eV.

should lie in the range of 0.75–1.0 and 1.0–1.25, respectively.⁵⁰ The compositional data obtained from energy-dispersive X-ray analysis (EDAX) are shown in Table 3. The EDAX shows elemental composition is near the stoichiometric ratio. But the increase in compositional ratios of Cu/(Zn + Sn) and (Zn/Sn) indicates the loss of Sn during sulfurization.

3.5. UV-vis spectroscopy analysis

The optical properties of CZTS films (optical absorption and bandgap) were studied using UV-visible spectroscopy. The absorbance plot of multi-stacked CZTS samples in Fig. 7(a) indicates that all precursors show high optical absorbance in the visible region. However, the 9LYS precursor shows significantly enhanced absorption in the visible region due to its high crystalline volume fraction and well-defined grain morphology. According to Tauc *et al.*,⁵¹ the optical band gap (E_g) calculations

were carried out. The region of the solar spectrum that a CZTS absorber will absorb depends on the optical band gap. By extrapolating the linear portion of the $(\alpha h\nu)^2 vs. h\nu$ plot shown in Fig. 7(b), it is possible to determine the band gap energy (E_g) for

Table 3 Energy dispersive X-ray analysis (EDAX) analysis of 3LYS, 6LYS, and 9LYS multi-stacked layers of CZTS films

	Elemental composition (atomic %)				Composition ratio		
Sample	Cu	Zn	Sn	S	Cu/ (Zn + Sn)	Zn/Sn	S/(Cu + Zn + Sn)
3LYS 6LYS 9LYS	21.12 23.37 24.03	12.61 14.27 15.09	11.14 12.35 11.43	55.13 50.01 49.45	0.88 0.87 0.90	1.13 1.15 1.32	1.22 1.01 0.98



Fig. 6 FE-SEM micrographs of multi-stacked (a) 3LYS, (b) 6YS, and (c) 9LYS CZTS thin films.



Fig. 7 (a) Absorption spectra of multi-stacked precursor CZTS thin films and (b) Tauc plot used to estimate optical energy band gap.

direct transitions. The band gap energies of all CZTS samples are observed in the range between 1.47 eV and 1.51 eV. Band gap variation is mainly attributed to variations in the homogeneity and crystallinity of the films, which in turn cause variations in the size of the crystallites between the samples.⁵² Therefore, the band gap values are relatively near the ideal band gap for solar cells.⁵³

3.6. Photoresponse and impedance measurements

Measurements of photoresponse and impedance were performed using a standard three-electrode system. For the measurements, a buffer layer of cadmium sulfide (CdS) was coated onto the multi-stacked CZTS thin films by sputtering. To maintain the spectral response of the multi-stacked precursors CZTS films, the thickness of the CdS layer was kept between 50– 60 nm. The schematic of prepared CZTS(3LYS/6LYS/9LYS)/CdS films for photoresponse and impedance measurement is shown in Fig. 8(a). Saturated calomel electrodes (SCE), platinum mesh, and CZTS(3LYS/6LYS/9LYS)/CdS on Mo-coated glass substrate were utilized as the working electrode, counter electrode and reference electrode, respectively. The schematic of the conventional three-electrode cell is shown in Fig. 8(b). Additionally, a 5 mM redox electrolyte of europium nitrate hydrate $[Eu(m)(NO_3)_3 \cdot 6H_2O]$ was employed as an electron scavenger.

The impedance measurement is susceptible to light illumination. In impedance spectra, the arc radius describes the junction resistance between the semiconductor surface and the



Fig. 8 Schematic of (a) CZTS/CdS junction (b) conventional threeelectrode cell employed in the present study.

electrolyte. An equivalent electrical circuit replicates and fits impedance spectroscopy data from the Nyquist plot. The inset of Fig. 9 displays the electrical parts used in the corresponding circuit.

After fitting electrochemical impedance spectroscopy (EIS) data with the equivalent circuit, 9LYS thin film showed the lowest charge transfer resistance ($R_{\rm ct} = 154.14 \ \Omega$) compared to the conventional 3LYS thin film ($R_{\rm ct} = 240.08 \ \Omega$), which suggested that it can help to improved electrical conductivity. It could be responsible for achieving high photocurrent in solar cell applications. The photoresponse of fabricated CZTS films, as quantified by the transient photocurrent density (*J*-*t*) upon the light irradiation (1000 W m⁻²) in AM 1.5, is shown in Fig. 10.



Fig. 9 Impedance spectra of synthesized CZTS thin films using multistacked layer precursors. The inset contains the equivalent circuit diagram.



Fig. 10 Transient photocurrent density as a function of time for multistacked precursors CZTS films biased at -0.5 V.

When the light is turned ON and OFF under chopping conditions, a rapid rise and fall of photocurrent are observed. When the electrode is illuminated, the photogenerated holes migrate to the back electrode (Mo), and the photogenerated electrons migrate into the electrolyte, resulting in a dramatic increase in photocurrent. In addition, for measurements, the active area of each sample was kept constant at 10 mm \times 10 mm, which is determined by the masked residual area. Interestingly, the photocurrent density in 9LYS CZTS was about twice that of 6LYS and four times larger than of conventional 3LYS CZTS film for a given bias voltage of -0.5 V because of higher light absorption, reduced electron-hole recombination, and increased charge transfer at the interface. Our prepared multistacked precursors 9LYS CZTS film shows a maximum photocurrent density of 82 μ A cm⁻² is much superior to the other CZTS reported in the literature.13,54-57 These findings suggest that the multi-stacked-layer modification dramatically impacts the performance of the CZTS absorber from 3LYS to 9LYS.

Fig. 10 shows the dark current (I_d) corresponding to all prepared CZTS films (for 3LYS $I_d = 14.7 \mu$ A, for 6LYS $I_d = 16.1 \mu$ A and for 9LYS $I_d = 15.2 \mu$ A) are nearby same but on the higher side as compared to work done previously.^{23,33,57,58} Therefore, one can further work on suppressing the dark current along with the increase in the photocurrent, which is a prerequisite for photoresponse enhancement, which allows multi-stacked precursors CZTS films to perform even better. Courel *et al.*⁵⁹ carried out a detailed analysis of different loss mechanisms and their contribution to the dark current density to understand what limits the performance of the CZTS-based solar cells.

Fig. 11 shows rise and decay curves (exponential fitted) of multi-stacked precursors CZTS films in the presence of 1000 W m⁻² light irradiation in AM 1.5 at a bias of -0.5 V. The fitted curves confirm the exponential time dependency. The estimated rise time (decay time) for 3LYS and 6LYS is 28 ms (120 ms) and 15 ms (92 ms), respectively. The rise and decay time obtained for 6LYS is significantly low compared to the reported literature.^{13,33} Though we observed high rise time (294 ms) and decay time (1980 ms) for the 9LYS sample as compared to 3LYS and 6LYS, the sensitivity, responsivity, and detectivity of the 9LYS sample shown in Table 3 are significantly enhanced due to



Fig. 11 Exponential fitted rise and decay photoconductivity curve of multi-stacked precursors CZTS films.

superior inter-diffusion of metallic precursors improved crystal quality of CZTS.

Sensitivity (S) is defined as the relative increase in the current when a light source illuminates the film and is measured using relation,⁶⁰

$$S(\%) = \frac{I_{\rm ph}}{I_{\rm D}} \times 100 \tag{5}$$

where $I_{\rm ph}$ = Photo current = ($I_{\rm L} - I_{\rm D}$), $I_{\rm L}$ is the light current, and $I_{\rm D}$ is the dark current.

Responsivity (*R*) is the amount of photocurrent generated when a light source illuminates the films and is evaluated by the formula,⁶¹

$$R = \frac{I_{\rm ph}}{PA} \tag{6}$$

where, A = Illumination area and P = Illumination light intensity.

The detectivity of prepared CZTS thin films was calculated by the relation,⁶²

Table 4 Photo current measurement of multi-stacked precursors CZTS films

Sample	$I_{\rm D}$ (μ A)	$I_{\rm L}$ (μA)	I_{ph} (μA)	S (%)	$R (\text{mA W}^{-1})$	$D imes 10^9$ (Jones)
3LYS	14.7	34.1	19.4	131	0.194	0.89
6LYS	16.1	54.2	38.1	237	0.381	1.68
9LYS	15.2	97.1	81.9	539	0.819	3.71

$$D = \frac{R \times A^{\frac{1}{2}}}{\left(2 \times e \times I_{\mathrm{D}}\right)^{\frac{1}{2}}} \tag{7}$$

where R is the responsivity and e is the charge of an electron.

The sensitivity (*S*), responsivity (*R*), and detectivity (*D*) of all photodetectors are tabulated in Table 4.

The responsivity of CZTS films increases from 0.194 to 0.819 mA W^{-1} with an increase of multi-stacked precursor from 3 layers to 9 layers, which may be due to the decrease in resistivity and increase of mobility of the films. It is, therefore, possible to forecast that thicker metallic precursor stacks with higher metallic precursor numbers, like 12 LYS, may be feasible. However, based on this study, the 9 LYS should be close to the restricted values due to two critical factors diffusion length and thickness of the individual layer. At the same time, a device with poor performance will arise from the precursor's (mostly sulfur's) interaction with the Mo substrate at greater sulfurization temperatures and over a more extended period. Multistacked 9-layer with sulfurization at 550 °C for 10 min is the best approach to improve device performance.

4 Conclusion

In conclusion, we showed that three types of high-quality CZTS thin films might be produced by sulfurizing a multi-stacking Sn/ZnS/Cu metallic precursor via RF-magnetron sputtering in an environment of H₂S and argon (kept at room temperature and ambient pressure). On an increasing number of stacked layers, i. e. from 3LYS to 9LYS system, there is an enhancement in the crystal quality, surface morphology, and photoresponse of CZTS films. Furthermore, in the case of the 9LYS sample, the improved inter-diffusion within stacked metallic layers resulted in a significant enhancement in crystallinity and homogenous grain size distribution throughout the film. Finally, in this work, CZTS thin film is fabricated via a modified stacked metallic precursor sequence exhibited a larger photocurrent density in the 9LYS (\sim 82 μ A cm⁻²) sample as compared to 6LYS (\sim 38 μ A cm^{-2}) and 3LYS (~19 μ A cm^{-2}) samples. These results demonstrate that optimizing stacked metallic precursor sequences while fabricating CZTS thin films can play a vital role in improving the overall performance of the realized films for solar cell applications.

Author contribution

Subhash Pandharkar: methodology, formal analysis, investigation, data curation, writing-original draft. Yogesh Hase: methodology, conceptualization, validation, formal analysis, investigation. Shruti Shah: methodology, validation, formal analysis, investigation. Vidya Doiphode: methodology, validation, formal analysis, investigation. Ashish Waghmare: methodology, conceptualization, validation, investigation. Ashvini Punde: methodology, validation, formal analysis, investigation. Pratibha Shinde: conceptualization, validation, formal analysis, investigation. Swati Rahane: data curation, formal analysis, investigation. Bharat Bade: conceptualization, validation, formal analysis, investigation. Somnath Ladhane: data curation, formal analysis, investigation. Mohit Prasad: data curation, writing-review, and editing. Shashikant P. Patole: writingreview and editing. Sandesh Jadkar: visualization, writingreview, editing, supervision, funding acquisition.

Conflicts of interest

There are no conflicts to declare.

Acknowledgements

Subhash Pandharkar, Yogesh Hase, Shruti Shah, Vidya Doiphode, Ashish Waghmare, Pratibha Shinde, and Bharat Bade are thankful to the Ministry of New and Renewable Energy (MNRE), Government of India, for the financial support under the National Renewable Energy Fellowship (NREF) program. Ashvini Punde is thankful to the Mahatma Jyotiba Phule Research and Training Institute (MAHAJYOTI), Government of Maharashtra, for the Mahatma Jyotiba Phule Research Fellowship (MJPRF). Swati Rahane is thankful for the research fellowship to the Chhatrapati Shahu Maharaj Research, Training and Human Development Institute (SARTHI), Government of Maharashtra. Finally, Mohit Prasad and Sandesh Jadkar are grateful to the Indo-French Centre for the Promotion of Advanced Research-CEFIPRA, Department of Science and Technology, New Delhi, for special financial support. Shashikant P. Patole would like to thank Khalifa University for its financial support through the internal fund for high-quality publications.

References

- 1 K. S. Gour, V. Karade, P. Babar, J. Park, D. M. Lee, V. N. Singh and J. H. Kim, *Sol. RRL*, 2021, 5, 2000815.
- 2 M. P. Suryawanshi, U. V. Ghorpade, C. Y. Toe, U. P. Suryawanshi, M. He, D. Zhang, J. S. Jang, S. W. Shin, J. H. Kim, X. Hao and R. Amal, *Prog. Mater. Sci.*, 2023, **134**, 101073.
- 3 F. Liu, Q. Zeng, J. Li, X. Hao, A. Ho-Baillie, J. Tang and M. A. Green, *Mater. Today*, 2020, **41**, 120.
- 4 H. Katagiri, K. Jimbo, W. S. Maw, K. Oishi, M. Yamazaki, H. Araki and A. Takeuchi, *Thin Solid Films*, 2009, **51**7, 2455.
- 5 S. Ge, H. Xu, Y. Huang, S. K. Karunakaran, R. Hong, J. Li, Y. Mai, E. Gu, X. Lin and G. Yang, *Sol. RRL*, 2020, 4, 2000325.
- 6 J. Jianga, L. Zhangb, W. Wanga and R. Hong, *Ceram. Int.*, 2018, 44, 11597.
- 7 L. Sun, J. He, H. Kong, F. Yue, P. Yang and J. Chu, *Sol. Energy Mater. Sol. Cells*, 2011, **95**, 2907.

View Article Online RSC Advances

Paper

- 8 S. Giraldo, Z. Jehl, M. Placidi, V. Izquierdo-Roca, A. Pérez-Rodríguez and E. Saucedo, *Adv. Mater.*, 2019, 31, 1806692.
 9 H. Katagiri, *Thin Solid Films*, 2005, 480–481, 426.
- 10 C. Zeng, Y. Liang, L. Zeng, L. Zhang, J. Zhou, P. Huang and R. Hong, *Sol. Energy Mater. Sol. Cells*, 2019, **203**, 110167.
- 11 Y. Liang, C. Zeng, L. Zeng, G. Yan, Y. Yuan, X. Lin, H. Zhu, Y. Mai and R. Hong, *Sol. RRL*, 2022, **6**, 2101047.
- 12 S. Ge, H. Xu, S. N. Khan, W. Yang, R. Hong, Y. Mai, E. Gu, X. Lin and G. Yang, *Sol. RRL*, 2021, **5**, 2100598.
- 13 O. P. Singh, A. Sharma, K. S. Gour, S. Husale and V. N. Singh, *Sol. Energy Mater. Sol. Cells*, 2016, **157**, 28.
- 14 F. Jiang, S. Ikeda, T. Harada and M. Matsumura, *Adv. Energy Mater.*, 2014, 4, 1301381.
- 15 Q. Guo, G. M. Ford, W. C. Yang, B. C. Walker, E. A. Stach, H. W. Hillhouse and R. Agrawal, *J. Am. Chem. Soc.*, 2010, 132, 17384.
- 16 W. Wang, M. T. Winkler, O. Gunawan, T. Gokmen, T. K. Todorov, Y. Zhu and D. B. Mitzi, *Adv. Energy Mater.*, 2014, 4, 1301465.
- 17 S. W. Shin, S. M. Pawar, C. Y. Park, J. H. Yun, J. H. Moon, J. H. Kim and J. Y. Lee, *Sol. Energy Mater. Sol. Cells*, 2011, 95, 3202.
- 18 G. Zoppi, I. Forbes, R. W. Miles, P. J. Dale, J. J. Scragg and L. M. Peter, *Prog. Photovolt.: Res. Appl.*, 2009, 17, 315.
- 19 C. Shi, G. Shi, Z. Chen, P. Yang and M. Yao, *Mater. Lett.*, 2012, 73, 89.
- 20 V. Tunuguntla, W. C. Chen, P. H. Shih, I. Shown, Y. R. Lin, J. S. Hwang, C. H. Lee, L. C. Chen and K. H. Chen, *J. Mater. Chem. A*, 2015, 3, 15324.
- 21 B. R. Bade, S. R. Rondiya, Y. A. Jadhav, M. M. Kamble, S. V. Barma, S. B. Jathar, M. P. Nasane, S. R. Jadkar, A. M. Funde and N. Y. Dzade, *J. Alloys Compd.*, 2021, 854, 157093.
- 22 S. C. Riha, B. A. Parkinson and A. L. Prieto, *J. Am. Chem. Soc.*, 2009, **131**, 12054.
- 23 L. S. Khanzada, M. A. Makhdoom, X. Lin, H. Azimi and C. J. Brabec, *Optik*, 2023, **272**, 170381.
- 24 M. A. Makhdoom, V. Sgobba, L. S. Khanzada, S. Fladischer and E. Spiecker, *Phys. Status Solidi A*, 2019, **2016**, 1900040.
- 25 S. Chen, A. Walsh, X. Gong and S. Wei, *Adv. Mater.*, 2013, 25, 1522.
- 26 D. H. Son, D. H. Kim, S. N. Park, K. J. Yang, D. Nam, H. Cheong and J. K. Kang, *Chem. Mater.*, 2015, **27**, 5180.
- 27 J. J. Scragg, T. Ericson, X. Fontané, V. Izquierdo-Roca, A. Pérez-Rodríguez, T. Kubart, M. Edoff and C. Platzer-Björkman, *Prog. Photovolt.: Res. Appl.*, 2014, 22, 10.
- 28 I. V. Dudchak and L. V. Piskach, *J. Alloys Compd.*, 2003, 351, 145.
- 29 S. Kim, J. A. Marquez, T. Unold and A. Walsh, *Energy Environ. Sci.*, 2020, **13**, 1481.
- 30 T. K. Todorov, J. Tang, S. Bag, O. Gunawan, T. Gokmen, Y. Zhu and D. B. Mitzi, *Adv. Energy Mater.*, 2013, 3, 34.
- 31 D. Shin, B. Saparov and D. B. Mitzi, *Adv. Energy Mater.*, 2017, 7, 1602366.
- 32 G. Yan, C. Zeng, Y. Yuan, G. Wang, G. Cen, L. Zeng, L. Zhang, Y. Fu, C. Zhao, R. Hong and W. Mai, ACS Appl. Mater. Interfaces, 2019, 11, 32097.

- 33 K. S. Gour, B. Bhattacharyya, O. P. Singh, A. K. Yadav, S. Husale and V. N. Singh, *J. Alloys Compd.*, 2018, 735, 285.
- 34 Y. Xie, C. Zhang, F. Yue, Y. Zhang, Y. Shic and T. Ma, *RSC Adv.*, 2013, **3**, 23264.
- 35 C. Zha, C. Ji, J. Zhang, L. Shen, X. Zhang, S. Dongb and N. Bao, *RSC Adv.*, 2016, 6, 107151.
- 36 A. Fairbrother, L. Fourdrinier, X. Fontané, V. Izquierdo-Roca,
 M. Dimitrievska, A. Pérez-Rodríguez and E. Saucedo, *J. Phys. Chem. C*, 2014, 118, 17291.
- 37 R. B. V. Chalapathy, G. S. Jung and B. T. Ahn, Sol. Energy Mater. Sol. Cells, 2011, 95, 3216.
- 38 W. C. Chen, C. Y. Chen, V. Tunuguntla, S. H. Lu, C. Su, C. H. Lee, K. H. Chen and L. C. Chen, *Nano Energy*, 2016, 30, 762.
- 39 E. A. Brandes, E. A. Brooks and G. B. Smithells, *Metal Reference Book*, Oxford, 1992.
- 40 S. M. Pandharkar, S. R. Rondiya, A. V. Rokade, B. B. Gabhale, H. M. Pathan and S. R. Jadkar, *Front. Mater.*, 2018, 5, 13.
- 41 B. D. Cullity and S. R. Stock, *Elements of X-Ray Diffraction*, Princeton Hall, 2001.
- 42 G. B. Harris, London, Edinburgh Dublin Philos. Mag. J. Sci., 1952, 43, 113.
- 43 V. T. Tiong, T. Hreid, G. Will, J. Bell and H. Wang, *Sci. Adv. Mater.*, 2014, **6**, 1467.
- 44 P. Prabeesh, I. P. Selvam and S. N. Potty, *Appl. Phys. A*, 2018, 124, 225.
- 45 A. Fairbrother, X. Fontané, V. Izquierdo-Roca, M. Espíndola-Rodríguez, S. López-Marino, M. Placidi, L. Calvo-Barrio, A. Pérez-Rodríguez and E. Saucedo, *Sol. Energy Mater. Sol. Cells*, 2013, 112, 97.
- 46 A. J. Cheng, M. Manno, A. Khare, C. Leighton, S. A. Campbell and E. S. Aydil, *J. Vac. Sci. Technol.*, *A*, 2011, **29**, 051203.
- 47 G. Y. Kim, J. R. Kim, W. Jo, K. D. Lee, J. Y. Kim, T. T. T. Nguyen and S. Yoon, *Curr. Appl. Phys.*, 2014, 14, 1665.
- 48 J. Xu, X. Yang, Q. D. Yang, T. L. Wong and C. S. Lee, *J. Phys. Chem. C*, 2012, **116**, 19718.
- 49 N. J. Choudhari, Y. Raviprakash, B. J. Fernandes and N. K. Udayashankar, *J. Alloys Compd.*, 2019, **799**, 314.
- 50 H. Chen, Q. Ye, X. He, J. Ding, Y. Zhang, J. Han, J. Liu, C. Liao, J. Meia and W. Lau, *Green Chem.*, 2014, 16, 3841.
- 51 J. Tauc and A. Menth, *J. Non. Cryst. Solids*, 1972, **8–10**, 569–585.
- 52 M. C. Rao and S. S. Basha, Results Phys., 2018, 9, 996.
- 53 S. M. Pawar, B. S. Pawar, A. V. Moholkar, D. S. Choi, J. H. Yun, J. H. Moon, S. S. Kolekar and J. H. Kim, *Electrochim. Acta*, 2010, 55, 4057.
- 54 R. Mudike, A. Bheemaraju, T. Rasheed, N. Singh, S. R. Dhage, P. D. Shivaramu and D. Rangappa, *Ceram. Int.*, 2022, **48**, 35666.
- 55 V. Mahalakshmi, D. Venugopal, K. Ramachandran and R. Ramesh, *J. Mater. Sci.: Mater. Electron.*, 2022, **33**, 8493.
- 56 Nisika, K. Kaur, K. Arora, A. H. Chowdhury, B. Bahrami, Q. Qiao and M. Kumar, *J. Appl. Phys.*, 2019, **126**, 193104.
- 57 Z. Hou, Y. Li, J. Liu, H. Shena and X. Huo, *New J. Chem.*, 2021, 45, 1743.
- 58 Y. F. Tay, S. S. Hadke, M. Zhang, N. Lim, S. Y. Chiamd and L. H. Wong, J. Mater. Chem. A, 2020, 8, 8862.

- 59 M. Courel, J. A. Andrade-Arvizu and O. Vigil-Galán, *Solid-State Electron.*, 2015, **111**, 243.
- 60 M. S. Mahdi, K. Ibrahim, N. M. Ahmed, A. Hmood,
 F. I. Mustafa, S. A. Azzez and M. Bououdina, *J. Alloys Compd.*, 2018, 735, 2256.
- 61 A. Sharma, B. Bhattacharyya, A. K. Srivastava, T. D. Senguttuvan and S. Husale, *Sci. Rep.*, 2016, **6**, 19138.
- 62 D.-H. Kwak, D.-H. Lim, H.-S. Ra, P. Ramasamya and J.-S. Lee, *RSC Adv.*, 2016, **6**, 65252.